23-26th November 2021, IMPT, Lyon, France





erc

European Research Counci Established by the European Commission

Detection, dynamics and impact of landslides

in the second of

Anne Mangeney^{1,1}

with specialists in seismology (C. Hibert, E. Stutzmann, Y. Capdeville, C. Levy, J. P. Montagner, etc.), mathematics (F. Bouchut, E. Fernandez-Nieto, G. Narbona-Reina, J. Sainte-Marie, J. Garres, J. Delgado-Sanchez), acoustics (J. De Rosny, X. Jia, R. Toussaint), etc.

¹ Institut de Physique du Globe de Paris, Université de Paris ¹ Equipe ANGE INRIA-Laboratoire Jacques Louis Lions

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Background

Master 1 in theoretical Physics, Paris VI Master 2, Oceanography, Meteorology, Environment, Paris VI

• PhD

Modelling of anisotropic ice flow in ice-sheets (LGGE, Grenoble)

• Post-doc

Numerical simulation of tsunami waves (CEA, Bruyeres-le-Chatel)

Present situation

Full professor, IPGP, University of Paris, **Seismology Group Associate researcher**, Jacques-Louis Lions Lab., Paris (2021-2017), Member of Institut Universitaire de France (since 2021) Head of the Master Natural Hazards, IPGP President of the CNRS section 18 (Earth and Telluric Planets) Member of the Natural Hazard Prevention committee (Minister)

International connections

Sabbatical year Inst. for Nonlinear Science, UC San Diego, 2006 Visiting Professor Seismolab, Caltech, USA, 2012

• Achievements 4 children







Trois missions de l'IPGP : recherche, observation des phénomènes naturels, enseignement et diffusion des connaissances en sciences de la Terre, des planètes et de l'environnement



Le projet de tutelle secondaire IRD concerne l'UMR 7154 mais l'IPGP est organisé de telle sorte que tous les membres de l'UMR ont accès aux moyens de l'UMS 3454 et aux activités d'enseignement. Le directeur du grand établissement IPGP est directeur de l'UMR 7154, de l'UMS 3454 et des Formations (avec un directeur adjoint pour chaque entité)





Organisation

- 1 UMR unique (IPGP, UP, CNRS, IGN, U Réunion)
- 1 UMS (IPGP, CNRS)
- 17 équipes de recherche (2 aux grands moulins)

Répartition des personnels permanents

• IPGP	43,2 %
• CNRS	29,3 %
 Université de Paris 	19,4 %
• IGN	4,8 %
Université Réunion	3,4 %

Budget consolidé 2020 : 45 M€ (Subvention pour charge de service public : 3,7 M€)

Formations

• 180 L, 125 M, 100 PhD

 licence ST, double licence ASTER, bachelor ScPo-IPGP UP, CPES Janson de Sailly-IPGP/UP, Licence pro gestion et traitement des déchets, licence pro méthodes d'exploration

et géophysique appliquée

- master ST, master Génie de l'environnement et industrie
- EUR Earth-Planets-Universe dans Smarts'UP (+ Labex UnivEarthS)





5/22



4 grandes thématiques de recherche qui regroupent les travaux des 17 équipes de recherche et des observatoires







- I Introduction : Natural landslides and their simulation
 - Ia Geophysics and mathematics
 - Ib Complexity and variability of natural landslides
 - Ic Examples of landslide simulations for hazard assessment
 - Id Challenges in geophysics and landslide modelling

II – Thin layer depth-averaged models for field-scale simulation

- IIa Thin layer approximation for granular flows
- IIb Accounting for complex topography
- IIc Need of two different reference frames for landslide tsunamis
- IId Application to real landslides : unexplained high mobility

III – Back to lab-scale experiments and simulation

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- IIIb μ (I) rheology in 2D models, comparison with other models (DEM, thin-layer)
- Illc Static-flowing interface (erosion/deposition)
- IIId Grain-fluid mixture model

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Key elements to address geophysical problems



Strong need of mathematical modelling, numerical methods, data analysis, signal processing, Machine Learning, statistics, surrogate models, uncertainty quantification, data assimilation, etc.

Basic ingredients for geophysics-mathematics link

Specialists in geophysics

Specialists in mathematics

Mathematical challenges will naturally occur

Solve 'geophysical' problems **WITH** geophysicists

I will not give any recipies or equations to solve, etc. but

I will illustrate this type of geophisics/mathematics interaction



'mathematical' concessions



Suropean Research Council stablished by the European Commission

Synergy submitted

You have to create your own challenge by taking risks due to the necessity of:

- accounting for the natural complexity of the processes and
 - simplifying the problem enough to solve it numerically

Make things as simple as possible, but not simpler. A. Einstein

Natural processes described by these equations

Mass and momentum conservation:

$$\frac{\partial \rho}{\partial t} + \nabla .(\rho \mathbf{u}) = 0 \qquad \mathbf{f} = \rho \mathbf{g}$$

$$\frac{\partial (\rho \mathbf{u})}{\partial t} = -\mathbf{u} \nabla .(\rho \mathbf{u}) - \rho \mathbf{u} . \nabla \mathbf{u} + \operatorname{div} \boldsymbol{\sigma} + \mathbf{f}$$

with different initial and boundary considitons and rheological laws

- Natural flows are rich et complex, strong uncertainties on the rheology and associated parameters
- The flows strongly depend on the complex boundary conditions (topography, etc.)
- Approximations based on physical considerations have to be performed for application to real processes....

For hazard assessment



The 2 November 2021 landslide at Mallama in Colombia, which killed 17 people. Image via the <u>Gobernacion de Nariño</u>. (from Landslides Blog D. Petley)

For hazard assessment



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Landslides

- Erosion processes at the Earth surface and on telluric planets
- Risk assessment on Earth in relation with seismic, volcanic, climate forcing



Broad objectives

• Destabilization: understand and quantify landslide occurrence/properties in link with external forcing

detection, localization, characterization (volume, ...)





explain and quantify the high mobility of natural landslides ...

Lack of field measurements of their dynamics

Mechanics and modelling of complex granular flows





Large variety of natural flows





Volume scale : $m^3 \rightarrow 10^5 \text{ km}^3$ Time scale : second \rightarrow year \neq Sources, \neq Topographies



On other planets



Ganges Chasma landslide, Valles Marineris, Mars



Gullies, mega-dune of Russell crater, Mars

Very different rheological laws



Mostly dry granular flows





La Réunion



Douvre, France

« Dry » granular flows



Snow avalanche : role of air





Significant erosion

High variability of snow behavior including thermal effects

Pyroclastic flows: role of gas



Debris avalanches + pyroclastic flows

Soufrière Hills, Montserrat 1995-2014 Significant erosion





Debris flows : role of water





Debris avalanches may transform into debris flows by entering rivers

Debris flows



Debris flows, Alps



Strong role of water Significant erosion

Debris flows, Iceland

Mud flows : strong role of water and fine particles



(Photo H. Hubl, Vienna)

QUINDICI, Italy

very rapid to extremely rapid flow of saturated debris in a channel, involving significantly greater water content relative to the source material

Several possible classifications...

Table 1. Main types of geophysical flows and typical ranges of values of most relevant parameters.

Flow type	Settin ambie	g, nt fluid	Interstitial fluid	Particle size (m)	Particle density (kg m ⁻³)	Particle volume fraction	Volume (m ³)
Subaerial landslides, rockfalls, rock or debris avalanches	Subaerial, extraterrestrial ^a		Air, none, small water content	$10^{-3} - 10^{1}$	~2000–3000	~0.4–0.7	$\frac{10^{0} - 10^{10}}{10^{9} - 10^{13a}}$
Submarine landslides	Subaqueous		Water			_	$10^{0} - 10^{13}$
Turbidity currents	Subaqueous		Water	$10^{-4} - 10^{-1}$	~1500–2500	~0.001-0.1	$10^{6} - 10^{10}$
Snow avalanches (dense, powder ^c)	Subaerial		Air (water)	$10^{-4} - 10^{-1}$	~100–1000	~0.1–0.4 ^b ~0.001–0.01 ^c	$10^4 - 10^6$
Pyroclastic density currents (dense ^d , dilute ^e)	Subaerial, subaqueous, extraterrestrial		Volcanic gases, air	$10^{-6} - 10^{0}$	~500–3000	~0.1–0.5 ^d ~0.001–0.01 ^e	$10^4 - 10^8$
Debris flows, lahars	Subaerial, extraterrestrial		Water	$10^{-4} - 10^{0}$	~2000–3000	~0.2–0.8	$10^4 - 10^9$
Flow type	Velocity (m s ⁻¹)	Thickness (m)	Runout distance (km)				
Subaerial landslides, rockfalls, rock	$10^{-1} - 10^2$	$10^{-1} - 10^2$	$ \begin{array}{r} 10^{0} - 10^{1} \\ 10^{1} - 10^{2} \end{array} $	Cas Delement et al. 2017			
Submarine landslides		$10^{-1} - 10^{2}$	$10^{1} - 10^{2}$	for a review			
Turbidity currents	$10^{0} - 10^{1}$	$10^{1}-10^{2}$	$10^{1} - 10^{3}$				
Snow avalanches (dense, powder ^c)	$10^{1}-10^{2}$	$10^{0} - 10^{1}$	$10^{-1} - 10^{0}$				
Pyroclastic density currents (dense ^d , dilute ^e)	10^{0} -10 ^{1d} 10 ¹ -10 ^{2e}	$10^{0}-10^{2}$	$10^{0}-10^{2}$				
Debris flows, lahars	$10^{0} - 10^{1}$	$10^{0} - 10^{1}$	$10^{0} - 10^{2}$				

Flow classification depending on the different regimes

Dimensionless numbers

defined as the ratio of relaxation time under load to shear time

$$I = I_{\rm pi} = \sqrt{\rho_{\rm p}} \dot{\gamma}^2 d^2 / P_{\rm p}$$

$$I = I_{\mathrm{v}} = rac{\eta_{\mathrm{f}}\dot{\gamma}}{P_{\mathrm{p}}}$$

$$I = I_{\rm fi} = \sqrt{\rho_{\rm f} \dot{\gamma}^2 d^2 / P_{\rm p}}$$

Particle inertia dominated regime (inertial number)

Viscous resistance dominated regime (viscous number)

Fluid inertia dominated regimes

Savage number : ratio of the inertial grain shear stress to the weight of flowing layer per unit surface

Bagnold number: ratio of inertial grain shear stress to viscous shear stress

$$N_{\rm Sav} = \frac{\rho_{\rm p} \dot{\gamma}^2 d^2}{(\rho_{\rm p} - \rho_{\rm f})gH} \qquad N_{\rm Bag} = \frac{\phi \rho_{\rm p} \dot{\gamma}^2 d^2}{(1 - \phi)\eta_{\rm f} \dot{\gamma}} \qquad N_{\rm Bag} \sim I_{\rm pf}^2$$

Complex initial conditions : several sub-events



Bingham Canyon Mine Utah, Avril 2013 65 Mm³





Retrogressive landslides

J-B de Chabalier, OVSG-IPGP - 21-12-2009





19 novembre 2010

Guadeloupe

Complex initial conditions

DOLOMITI m. 1011 S.D ITO di CADURE BENVENUTI - BIENVENUS WELCOME - WILLKOMMEN

9th October 2021 – Dolomites, Italy

Induced hazards





Volcanic eruption











Volcanic flank collapse

History of volcanic domes: phases of construction/destruction



Volcanic flank collapse

12 July 2003 :

Soufrière Hills volcano, Montserrat (~200 Mm³) reaching the sea and generating a tsunami







All these processes are simulated in field conditions

Advanced numerical models are now used accounting for complex topographies, a free surface, complex initial conditions, complex rheological laws, etc.
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Beyond simple models



Past debris avalanche in Guadeloupe

First order role of:

• topography: channelling, overflows

Fine description of the topography: High numerical cost

• rheological behaviour: friction law, role of fluids, etc.

Very complex behaviour: Strong non-linearities

Peruzzetto et al., 2019, 2021

Beyond simple models





Link with seismology, volcanology, climatology, ocean and atmosphere...



snapshot of seismo-acoustic wave coupling





Kuehnert et al., 2020, 2021

Complex numerical models for hazard assessment



Figure 1 : Les nouveaux sites éruptif et sismiques de Mayotte et les pentes tsunamigènes du talus récifal (source REVOSIMA).

Tsunamis hazard due potential submarine landslides

Seismic activity + high submarine slopes



IPGP - BRGM - French Minister, local actors of hazards, Poulain et al., 2021

Potential landslide scenarios



Propagation of water waves



Arrival time and shape of the wave at the coast



Poulain et al., 2021

Numerical simulation of water surelevation



Maximum water detph during the simulation



Maximum water velocity during the simulation



Poulain et al., 2021

Hazard maps

BRGM-IPGP



Poulain et al., 2021

Evacuation plans and refuges for population



Leone et al., 2021

Evacuation plans and refuges for population



Leone et al., 2021

Evacuation plans and refuges for population



Leone et al., 2021

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Main questions in geophysics



- How to detect geophysical flows and to assess their related hazards and indirect impact (tsunamis, etc)?
- What is the contribution of gravitational flows in erosion processes and relief evolution at the surface of the Earth and other planets?
- How are gravitational flows related to external forcing? Could they provide indicators or precursors of these forcing processes?
- What physical processes may be at the origin of the high mobility of large landslides?
- How to quantify and model erosion/deposition processes, solid/fluid interaction, polydispersity and fragmentation at the natural scale?
- How to retrieve the mechanisms of propagation and the characteristics of the flows from their deposit and/or from the generated seismic or geophysical signal?

Review paper: Delannay, Valance, Mangeney, Roche, Richard, 2017

Current challenges for landslide modelling

Improve the models, beyond simple approaches, to :

- Better describe essential elements of natural rheological behavior and complexity (role of fluids, erosion/deposition processes, material properties, etc.)
- Decrease the computational cost by developping physically and mathematically relevant approximations for field-scale simulations but still describe complex topography effects
- Build stable numerical models

in order to :

- Simulate the flow dynamics and deposit and associated hazard
- Extract relevant information on landslides from field measurements

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Thin layer approximation



Role of topography, Earth sphericity, etc.

Shallow natural gravitational flows

rivers





Lava flows

Strong
thermal
effects

Glaciers



Shallow flows



Ice sheet flow

Numerical modelling of landslides

Natural materials







Modeling



Reasonable computational costHigh computationalHigh computationalEmpirical flow law λ costParticle size $\mu = tan \delta$ Local flow lawdistribution ???

Thin layer approximation on 2D topography

- Flow on complex natural topography high computational cost approximation: small aspect ratio $a = \frac{H}{L} \ll 1$
- Depth-averaged thin layer model





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Thin layer approximation on 3D topography

• Full curvature tensor

$$\mathcal{H} = c^3 \ \partial^2_{\mathbf{xx}} b$$

First equations including these « centrifugal » forces

SHALTOP X3 n × X2 \mathcal{R}_{x} h **X**1 Y, b X

Bouchut, Mangeney-Castelnau, Perthame, Vilotte, 2003, Bouchut and Westdickenberg 2004, Mangeney et al. 2007



SHALTOP equations in a Cartesian reference frame

• Momentum conservation equation in the Cartesian reference frame

$$\vec{\mathcal{V}} = \vec{V}_{1}\vec{e}_{X} + \vec{V}_{2}\vec{e}_{Y} + \frac{1}{c}\mathbf{s}^{t}\vec{\nabla}\vec{e}_{Z} \quad \text{with} \quad \vec{\nabla} = (\vec{V}_{1},\vec{V}_{2})$$

$$\partial_{t}\vec{\nabla} + (\vec{\nabla}\nabla_{\mathbf{x}})\vec{\nabla} + (I_{2} - \mathbf{ss}^{t})\nabla_{\mathbf{x}}\left(g(hc + b)\right)$$

$$= -c\left(\vec{\nabla}^{t}(\partial_{\mathbf{xx}}^{2}b)\vec{\nabla}\right)\mathbf{s} - \frac{\mu gc\vec{\nabla}}{\sqrt{\nabla}^{2}} + \left(\frac{1}{c}\mathbf{s}^{t}\vec{\nabla}\right)^{2} \left(1 + \frac{\vec{\nabla}^{t}(\partial_{\mathbf{xx}}^{2}b)\vec{\nabla}}{F_{\mu}}\right)_{+}$$

$$\vec{\nabla}^{2} + \left(\frac{1}{c}\mathbf{s}^{t}\vec{\nabla}\right)^{2} \left(1 + \frac{g}{F_{\mu}}\right)_{+} (Luca \ et \ al. \ (2009))$$

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Curvature effects on simplified topographies



Friction angle $\delta = 0^\circ$, slope angle $\theta = 10^\circ$

Strong effect for river or water flow simulations





Curvature effects on complex natural topographies



Model benchmark exercice



Gueugneau et al., 2021

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Two different integration directions for landslides and tsunami

For a review see Yavari-Ramshe & Ataie-Ashtiani (2016)



Delgado-Sanchez, Bouchut, Fernandez-Nieto, Mangeney, Narbona-Reina (2021)

Two different directions for landslides and tsunami

$$(X, t) \mapsto (x, t) = (X - (\tilde{b}(X) + H_2(X, t)) \sin \theta, t)$$

$$H_1(X, t) = h_1 (X - (\tilde{b}(X) + H_2(X, t)) \sin \theta, t),$$

$$U_1(X, t) = u_1 (X - (\tilde{b}(X) + H_2(X, t)) \sin \theta, t),$$

$$\mathcal{J} := \det(\nabla_{(X,t)}(x, t)) = 1 - \partial_X (\tilde{b} + H_2) \sin \theta$$

$$A := \nabla_{(X,t)}(x, t) = \begin{pmatrix} 1 - \partial_X (\tilde{b} + H_2) \sin \theta & -\partial_t H_2 \sin \theta \\ 0 & 1 \end{pmatrix}$$

$$\frac{\partial_t (H_1 \mathcal{J}) + \partial_X (H_1 (U_1 + \partial_t H_2 \sin \theta)) = 0,}{\partial_t (H_2 U_2) + \cos \theta_X (H_2 U_2) = 0,}$$
Friction between water and avalanche layers and avalanche

Numerical method to solve these equations

Essential step for application in geophysics !

$$W = (\mathcal{H}_1, \mathcal{Q}_1, H_2, Q_2)^T = (H_1 \mathcal{J}, H_1 \mathcal{J} U_1, H_2, H_2 U_2)^T$$

$$W_i^{n+1} = W_i^n - \frac{\Delta t}{\Delta x} \left(\tilde{\mathcal{F}}_{i+1/2}^n - \tilde{\mathcal{F}}_{i-1/2}^n \right) - \frac{\Delta t}{\Delta x} \mathcal{S}_i^n$$

$$\begin{split} \tilde{\mathcal{F}}_{i+1/2} &= \frac{1}{2} \Big(\mathcal{F}(W_{i+1/2}^+) + \mathcal{F}(W_{i+1/2}^-) \Big) \\ &- \frac{1}{2} \Big(\alpha_{0,i+1/2} \Big(W_{i+1/2}^+ - W_{i+1/2}^-) + \alpha_{1,i+1/2} \Big(\mathcal{F}(W_{i+1/2}^+) - \mathcal{F} \Big(W_{i+1/2}^-) + \mathcal{S}_{i+1/2} \Big) \Big) \end{split}$$

Hydrostatic reconstruction of the variables

$$W_{i+1/2}^{\pm} = (\mathcal{H}_{1,i+1/2}^{\pm}, \mathcal{Q}_{1,i+1/2}^{\pm}, H_{2,i+1/2}^{\pm}, Q_{2,i+1/2}^{\pm})^{T}$$

Quantification of the error in reference choice



Introduce curvature effects

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Simulation of natural flows

Simulation of observed deposits (Switzerland)

 $\mu = an \delta$: empirical description of the mean dissipation



t = 70 s



Calibrated friction angle : $\delta = 17^{\circ}$

Small compared to friction angles of natural materials ! $\theta_r \sim 35^{\circ}$

Origin of the high mobility of natural flows ?

Pirulli and Mangeney, 2008

Simulation of a large variety of natural flows

Simulation with empirical friction coefficient



Lucas, Mangeney, Ampuero, Nature Communications, 2014

Empirical friction laws based on deposit data



Friction coefficient μ = tan δ

Lucas, Mangeney, Ampuero, Nature Communications, 2014

Empirical friction weakening with velocity



Lucas, Mangeney, Ampuero, Nature Communications, 2014



Different physical processes



Fluid phase (water, mud)



Fluidization (gas, air)









// snow avalanche

// rivers, glaciers